

# Spec-Gloss Surfels and Normal-Diffuse Priors for Relightable Glossy Objects

Georgios Kouros, Minye Wu, Tinne Tuytelaars

Processing, Speech and Images (PSI) group

Department of Electrical Engineering (ESAT), KU Leuven, Belgium

# What is inverse Rendering?

Inverse Rendering

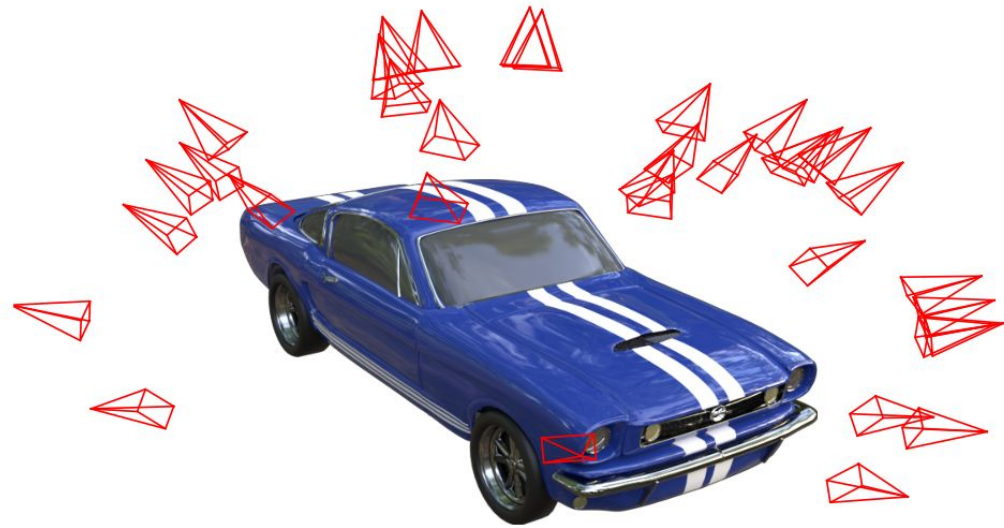


Learned 3D Scene  
Decomposition



Downstream Tasks

Training Views



albedo

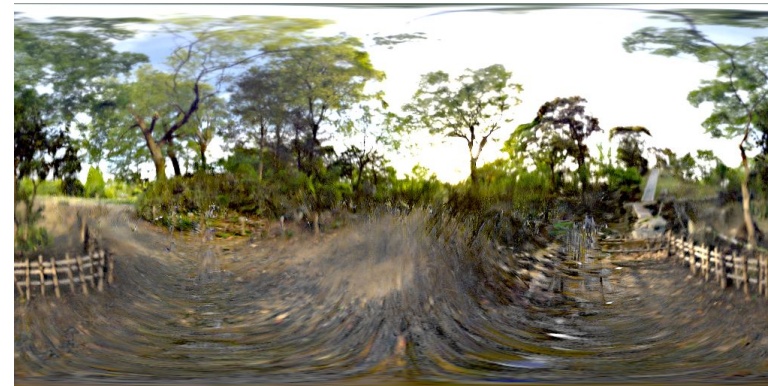
rough.

$F_0$

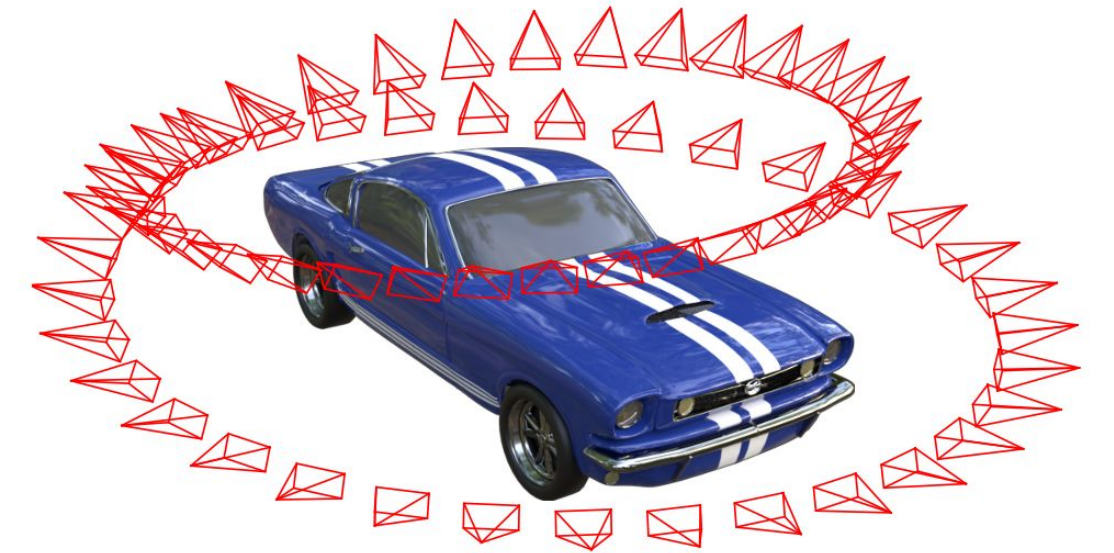
normals



Recovered Scene Illumination



Novel View Synthesis



Scene Editing



# Why does inverse rendering of glossy objects fail?

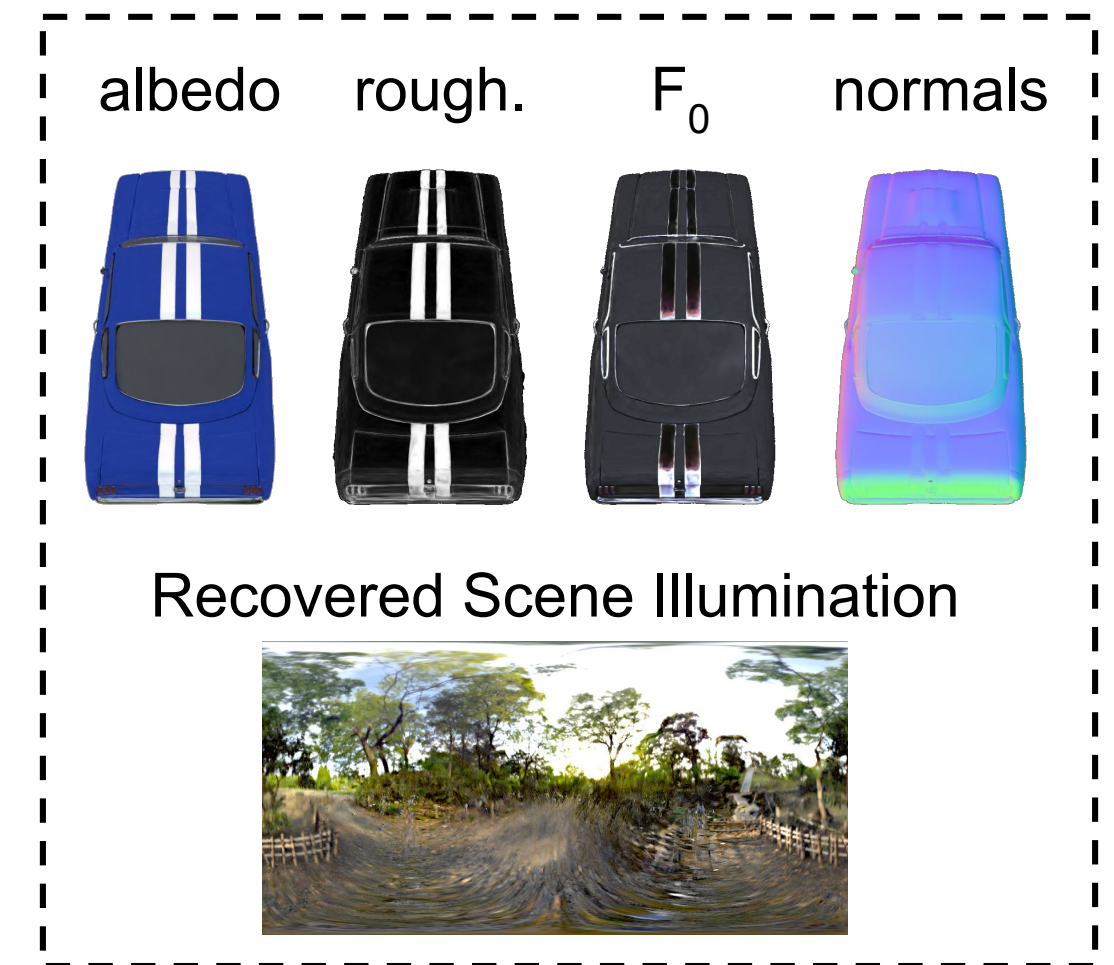
Inverse rendering is ill-posed because it jointly estimates

- **geometry** e.g. volumetric field, voxel grid, Gaussians
- **material properties** e.g. albedo, roughness,  $F_0$
- **scene illumination** e.g. HDR envmap

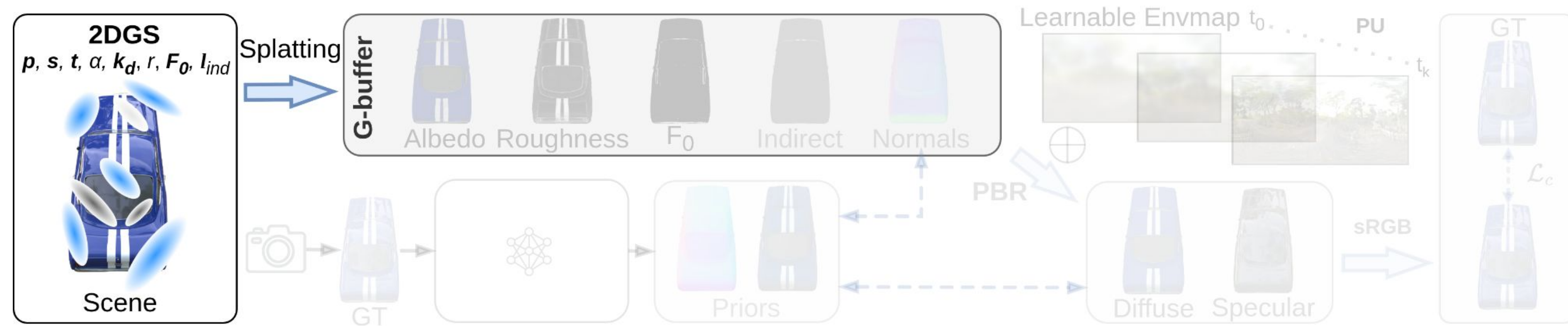
and is susceptible to local minima due to ambiguity from complex view-dependent effects.

This often leads to:

- **Good** novel view synthesis 😊, but
- **Bad** relighting / scene editing 😞

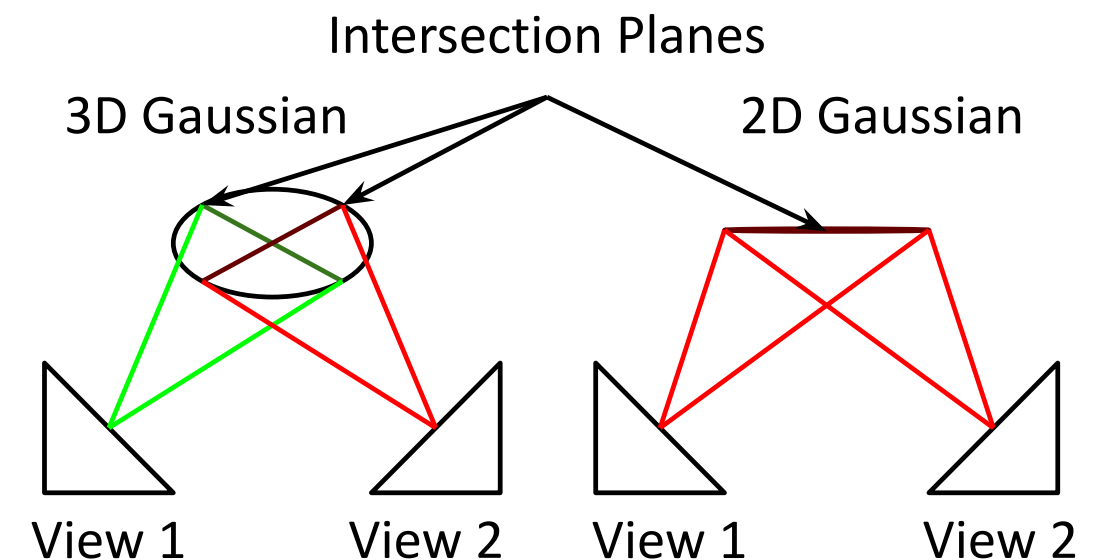


# Our pipeline (1/6): 2DGS + deferred shading



## Why **2D Gaussians** and not **3D Gaussians**?

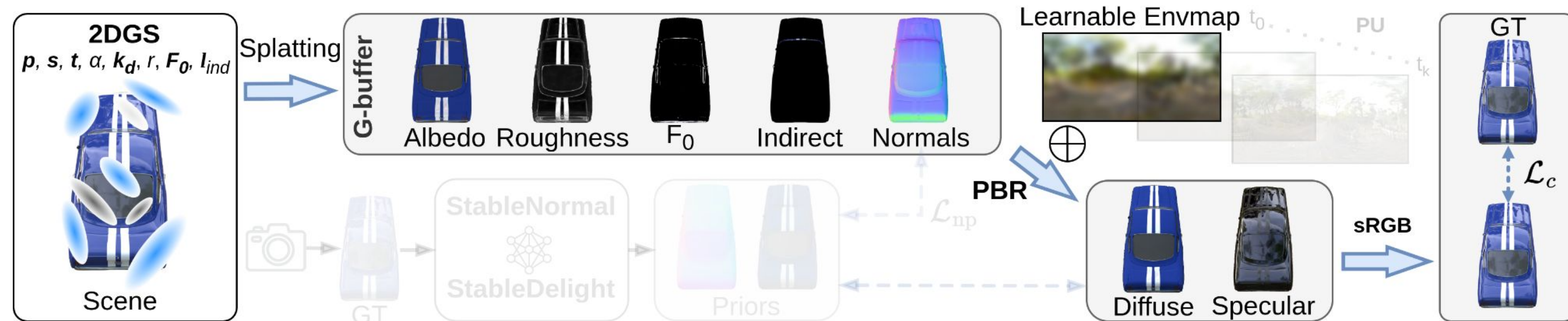
- View-consistent geometry
- Explicit normals as shortest axes of Gaussians



## Why **deferred shading** (pixel-level) and not **forward shading** (Gaussian level)?

- Improved reconstruction of high-frequency reflections with less blurring
- Improved scaling on large scenes (millions of Gaussians)

# Our pipeline (2/6): Spec-Gloss Material Parameterization



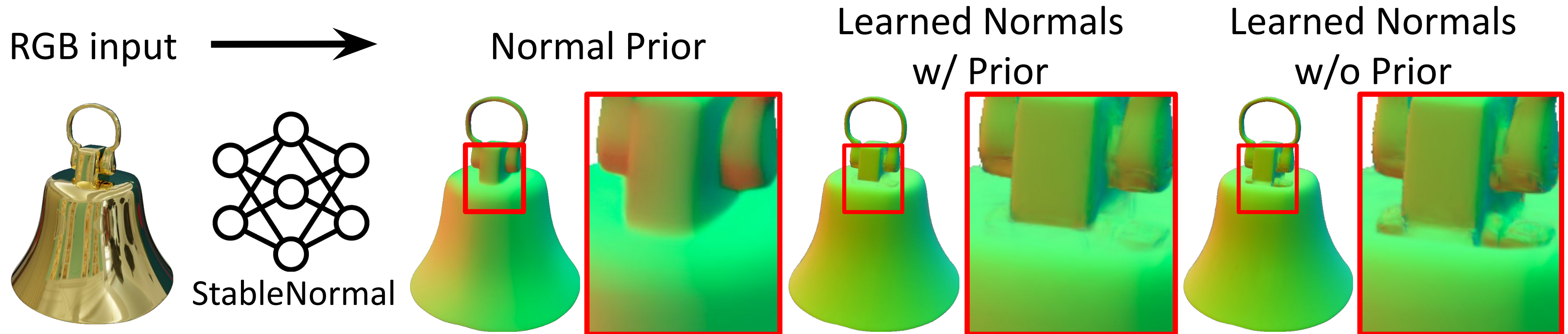
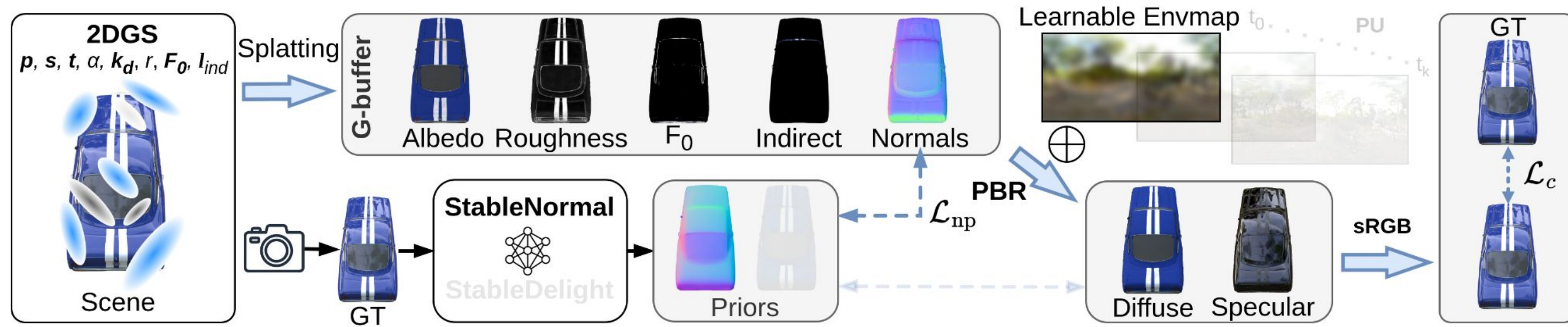
- The popular **Metallic-Roughness (MR)** parameterization entangles albedo / base-color ( $b$ ) in specular appearance via  $F_0$ :

$$F_0(m, \boxed{b}) = (1 - m) \cdot 0.04 + m \cdot \boxed{b}$$

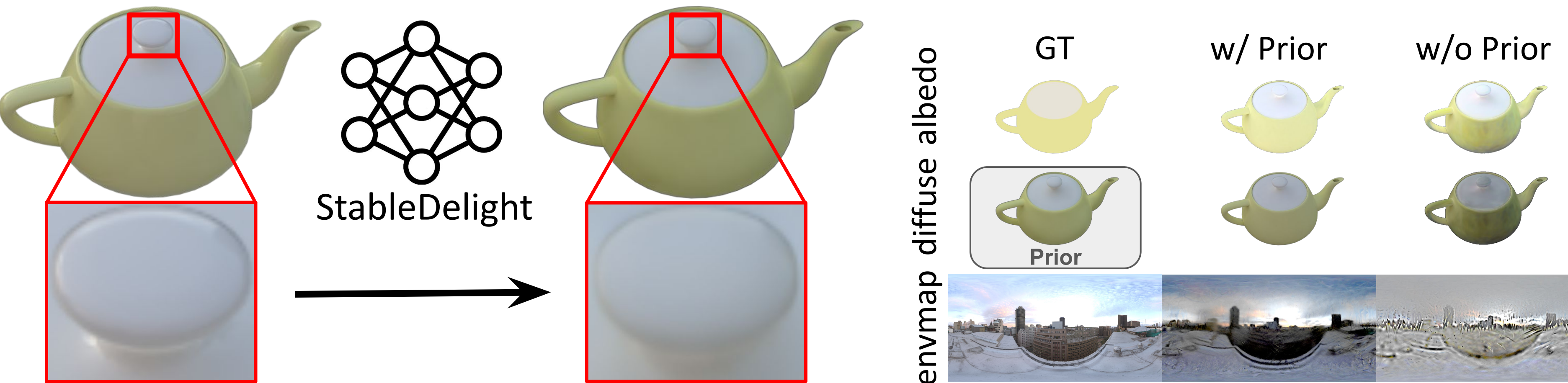
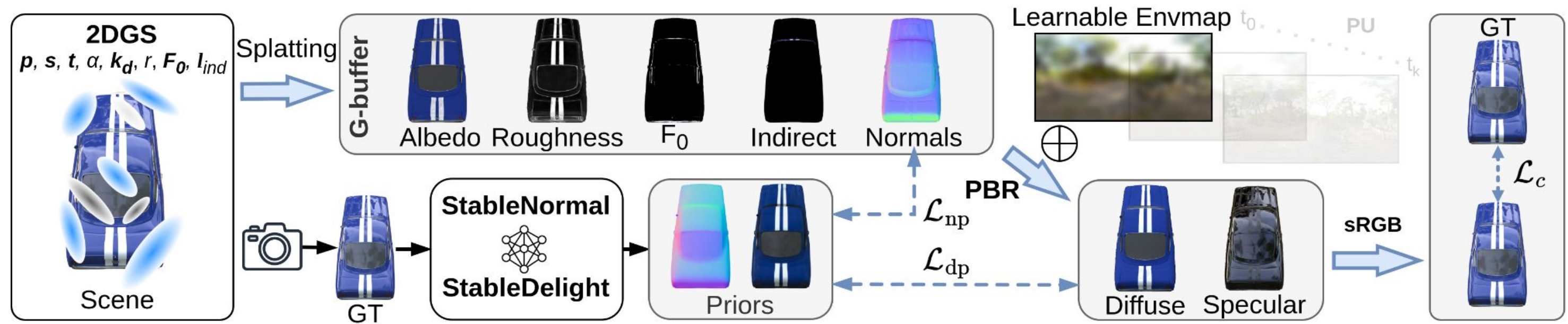
- The **Specular-Glossiness (SG)** parameterization learns  $F_0$  as a free parameter !!!



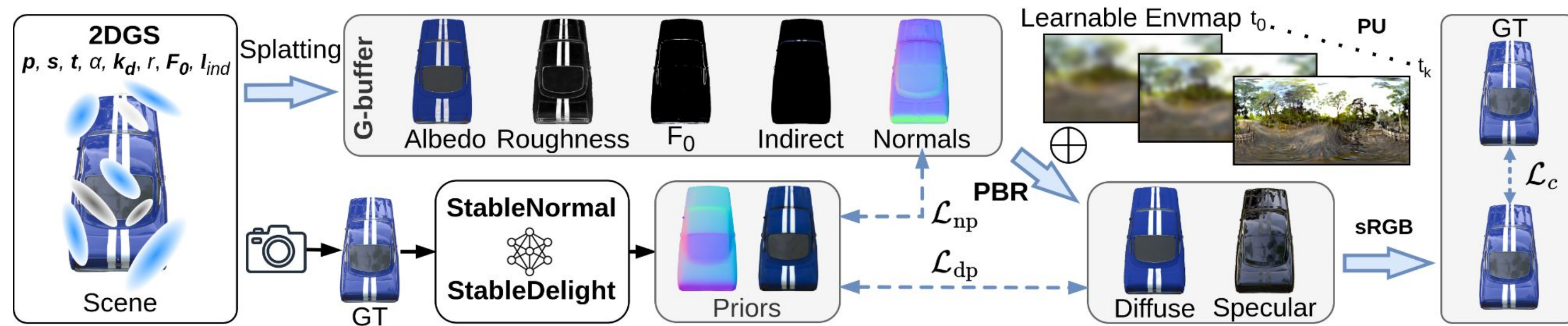
# Our pipeline (3/6): Surface normal prior



# Our pipeline (4/6): Diffuse color prior

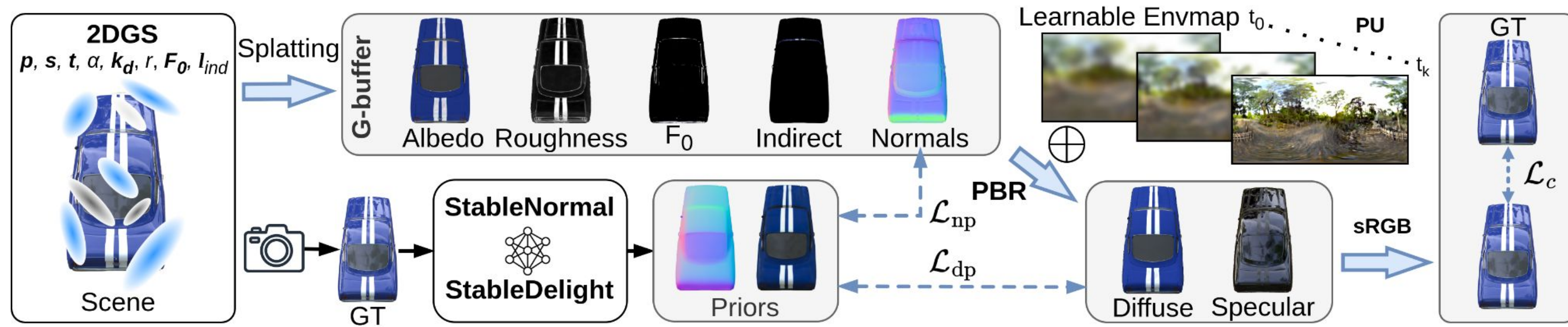


# Our pipeline (5/6): Coarse-to-fine lighting optimization



- **Envmap prefiltering** is the main bottleneck: higher res  $\rightarrow$  slower training
- Our method starts from a low resolution environment light cubemap ( $6 \times 64 \times 64$ ) and **progressively upsamples (PU)** it every 15k iters until the full resolution ( $6 \times 512 \times 512$ )
- **PU halves training time** (2h  $\rightarrow$  1h) without impacting envmap quality
- Facilitates recovery of **higher resolution envmaps** in reasonable time

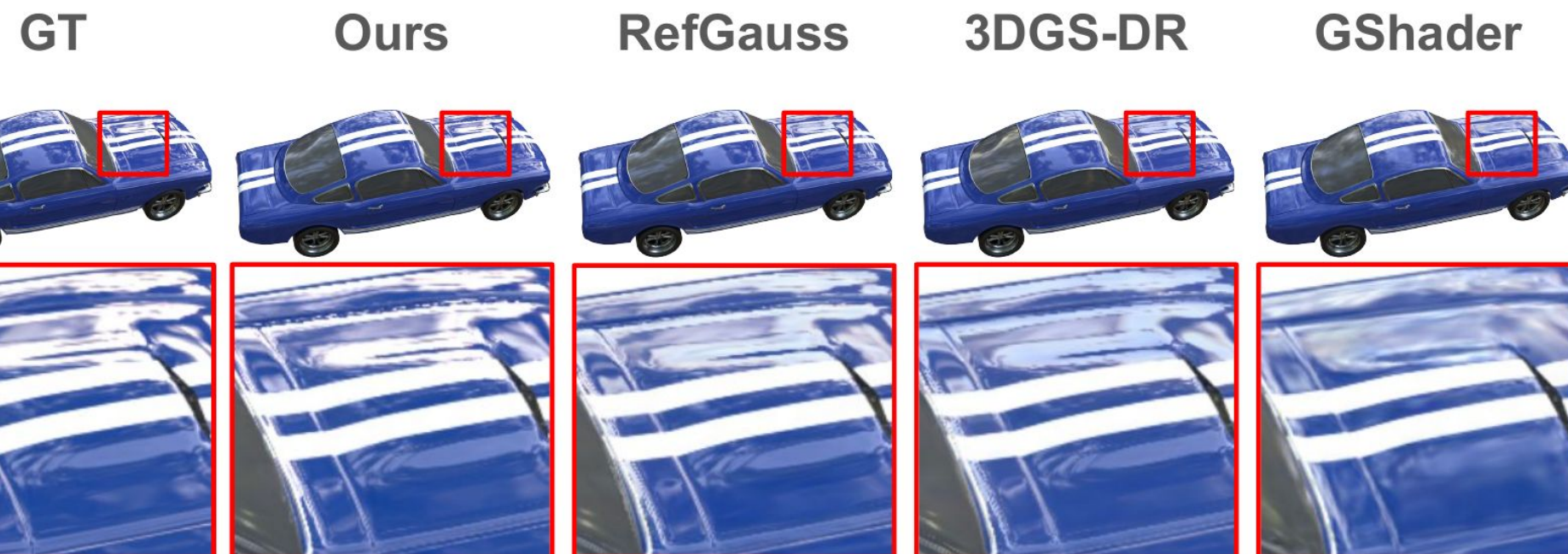
# Our pipeline (6/6): Negative-only Clipping (NOC)



$$E \leftarrow \max(0, E_{\text{raw}})$$

Preserves sharp bright specular reflections!

Improves albedo-lighting disentanglement!



# Experimental Results

# Envmap Recovery

Scene

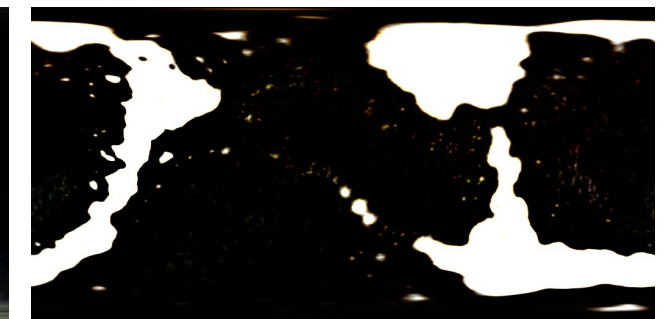
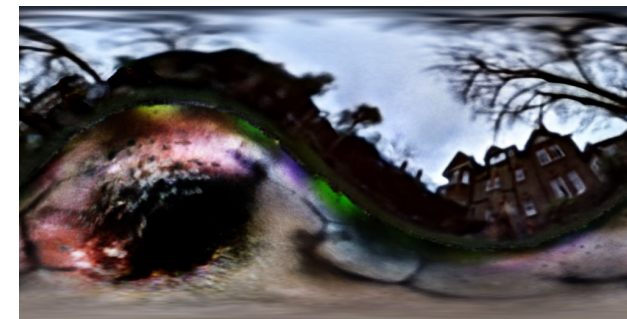
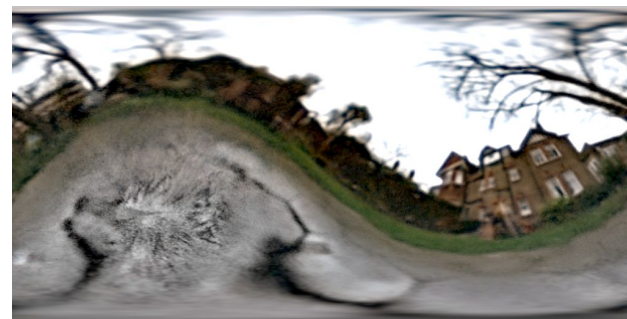
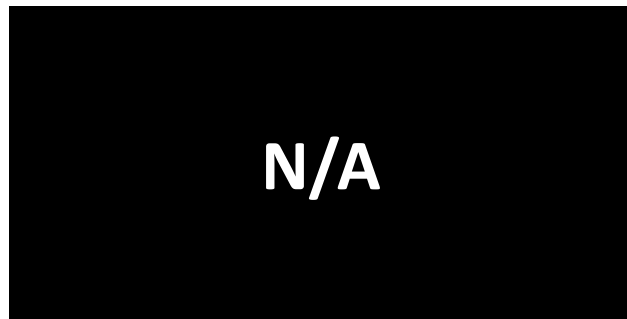
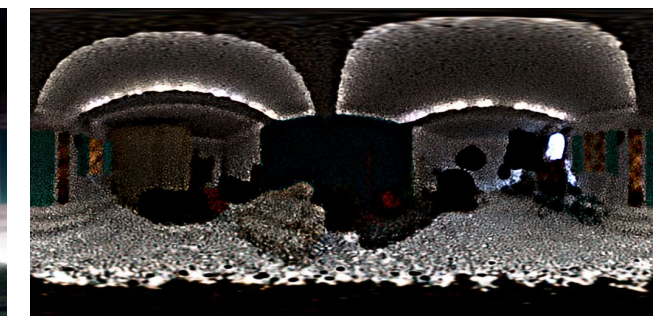
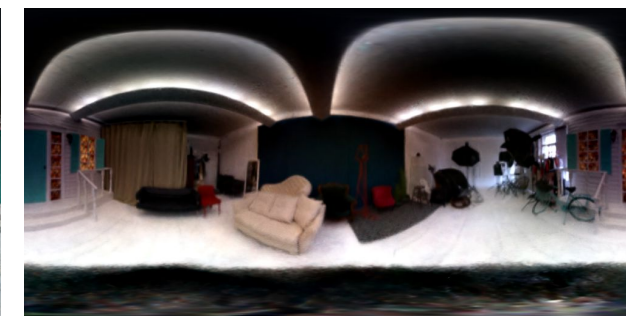
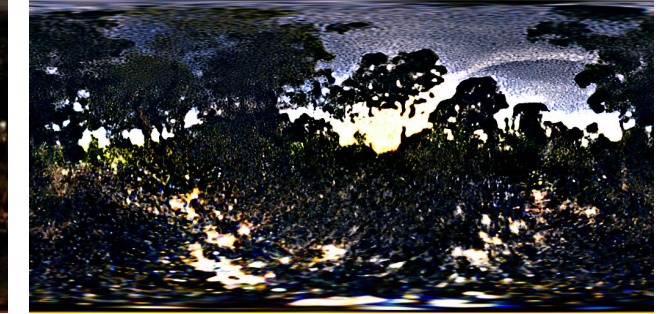
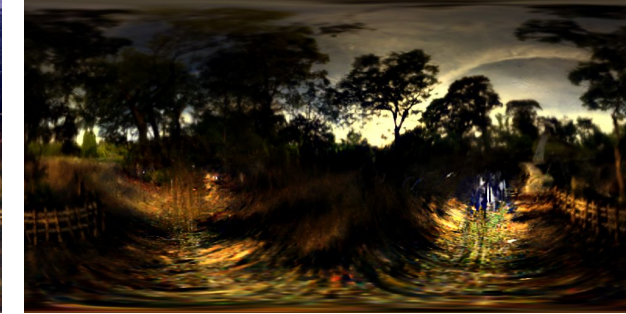
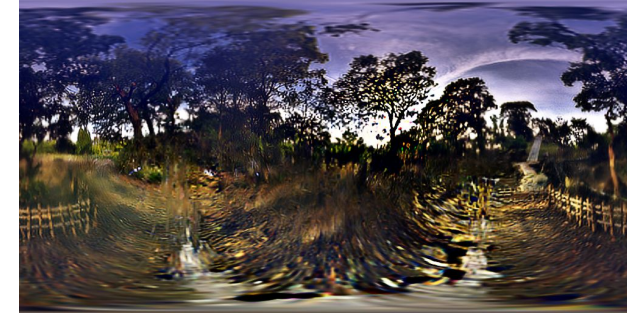
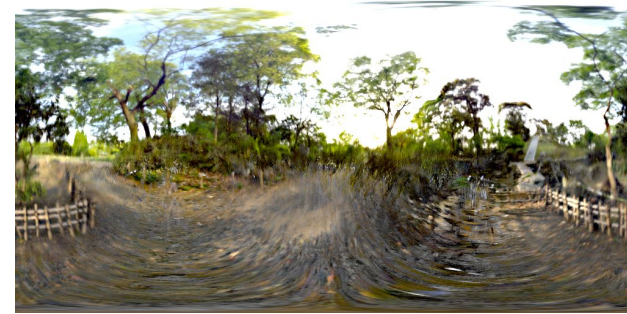
GT

Ours

Ref-Gaussian

3DGS-DR

GaussianShader



- Superior envmap recovery compared to other Gaussian Splatting methods!
- Accurate envmap → better material properties → high-quality relighting!

# Relighting Results

Relighting: bell

Ours (SG)

Ours (MR)

GT



Applied Envmap



# Novel View Synthesis Results

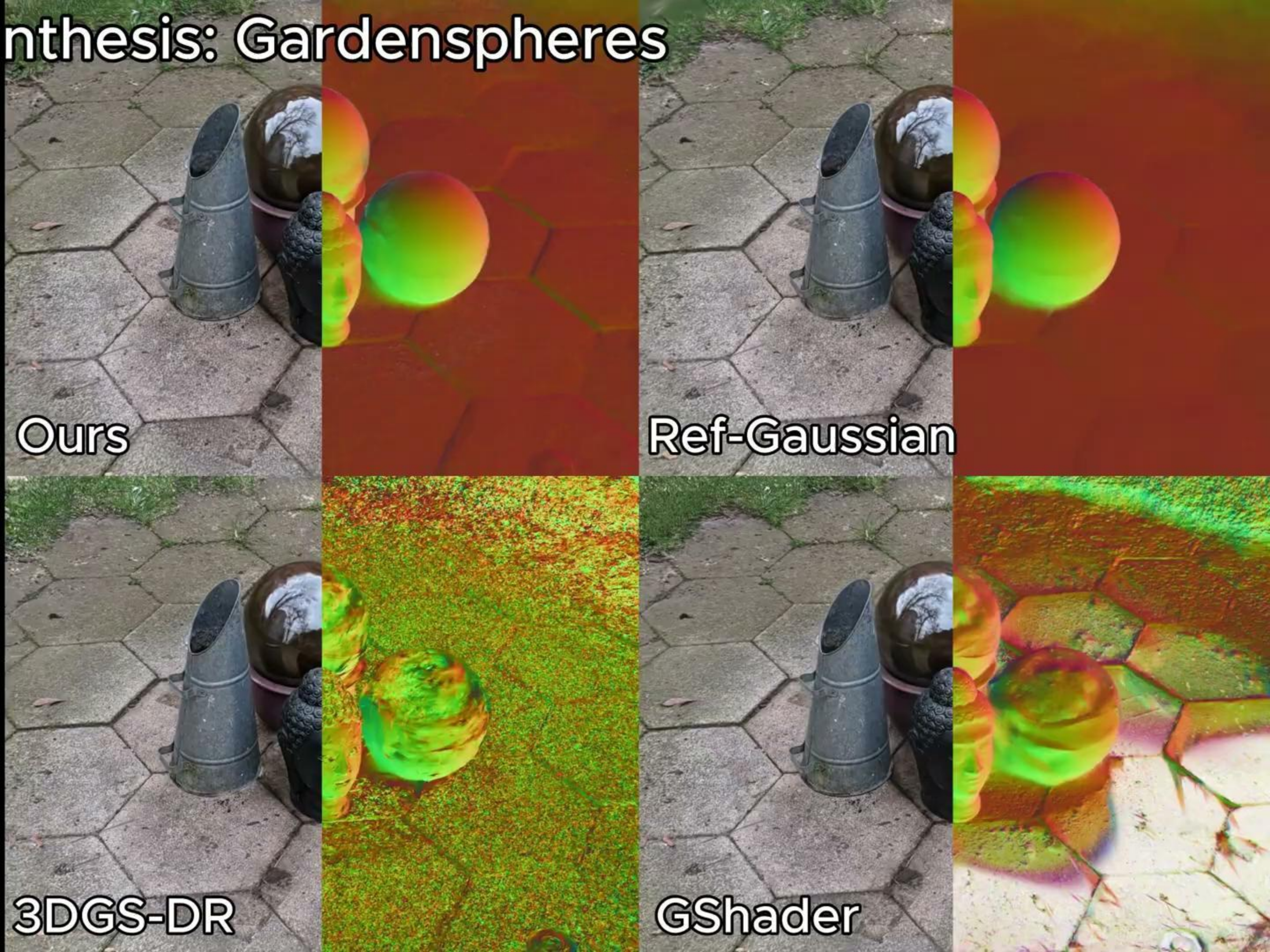
# Novel View Synthesis: Gardenspheres

Ours

Ref-Gaussian

3DGS-DR

GShader



# Novel View Synthesis: car



GT

Ours

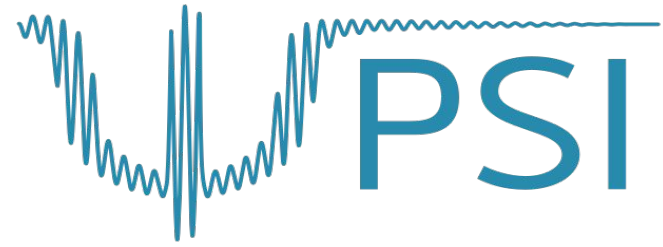
# Scene Editing Results

Original



Albedo Editing





Thank you for your attention!  
See you at **Poster Session 4-13!**

Project Page

